

Que: 1 (a)

Refer page 92 (Data book)

$$X = L/D = \frac{\text{Length side}}{\text{Distance}}$$

$$X = \frac{1}{0.5} = 2$$

$$Y = \frac{B}{D} = \frac{\text{Smaller side}}{\text{Distance}} = \frac{0.5}{0.5} = 1$$

from graph / Table $F_{1-2} = 0.285$

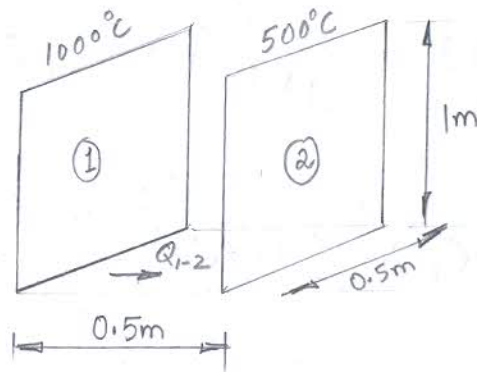
$$Q_{1-2} = A_1 F_{1-2} \sigma (T_1^4 - T_2^4) \quad \text{where } T_1 = 1000 + 273 = 1273 \text{ K}$$
$$T_2 = 500 + 273 = 773 \text{ K}$$

$$Q_{1-2} = (0.5 \times 1) (0.285) (5.67 \times 10^{-8}) (1273^4 - 773^4)$$

$$Q_{1-2} = 18333.54 \text{ W.}$$

Que: 1-b Wien's displacement law states that, the maximum ~~power~~ value of the emissive power E_b can be obtained by differentiating Planck's equation with respect to λ and equating it to zero, therefore:

(1)



$$\frac{dE_{b\lambda}}{d\lambda} = 0$$

$$\frac{dE_{b\lambda}}{d\lambda} = \frac{d}{d\lambda} \left[c_1 \lambda^{-5} \left(e^{\frac{c_2}{\lambda T}} - 1 \right)^{-1} \right] = 0$$

$$= c_1 \lambda^{-5} (-1) \left(e^{\frac{c_2}{\lambda T}} - 1 \right)^{-2} \left(-\frac{c_2}{\lambda T} \right) \left(e^{\frac{c_2}{\lambda T}} - 0 \right) + \left(e^{\frac{c_2}{\lambda T}} - 1 \right)^{-1} \cdot c_1 (-5) \lambda^{-6} = 0$$

$$= \left(c_1 \frac{1}{\lambda^6} \left(e^{\frac{c_2}{\lambda T}} - 1 \right)^{-2} \cdot \frac{c_2}{\lambda T} \left(e^{\frac{c_2}{\lambda T}} \right) \right) = 5 c_1 \left(e^{\frac{c_2}{\lambda T}} - 1 \right)^{-1} \frac{1}{\lambda^6} = 0$$

$$\frac{e^{\frac{c_2}{\lambda T}}}{\left(e^{\frac{c_2}{\lambda T}} - 1 \right)} = \frac{5\lambda T}{c_2}$$

put $\frac{c_2}{\lambda T} = x$

$$\frac{e^x}{e^x - 1} = \frac{5}{x} \Rightarrow$$

$$x e^x = 5 e^x - 5$$

$$x e^x - 5 e^x + 5 = 0$$

$$5 e^x \left[\frac{x}{5} - 1 + e^{-x} \right] = 0$$

$$e + \frac{x}{5} - 1 = 0$$

solving the above equation gives $x = 4.965 = \frac{c_2}{\lambda_{\max} \cdot T}$

$$\lambda_{\max} \cdot T = \frac{c_2}{4.965} = \frac{1.438 \times 10^{-2}}{4.965} = \boxed{2.898 \times 10^{-3} \text{ m-k}}$$

$$\lambda_{\max} \cdot T = 2.898 \times 10^{-3} \text{ m-k}$$

(2)

Que: 2

The amount of radiation directed towards the area dA_p

$$dE_b = I_n dA_1 \cos \phi \frac{dA}{r^2}$$

area of the small element

$$dA_p = (R \sin \theta, d\theta) (R \cdot d\phi)$$

$$= R^2 \sin \theta \cdot d\theta \cdot d\phi$$

solid angle in the direction of $\theta = d\omega = \frac{dA_p}{R^2}$

$$d\omega = \frac{dA_p}{R^2} = \frac{R^2 \sin \theta \cdot d\theta \cdot d\phi}{R^2} = \sin \theta \cdot d\theta \cdot d\phi$$

$$dE_b = I_n \cdot \cos \theta \cdot d\omega \cdot dA_1$$

$$= I_n \cdot \cos \theta \cdot \sin \theta \cdot d\theta \cdot d\phi \cdot dA_1$$

$$\int dE_b = I_n dA_1 \int_0^{\pi/2} \sin \theta \cdot \cos \theta \cdot d\theta \cdot \int_0^{2\pi} d\phi$$

$$E_b = I_n \cdot dA \cdot \frac{1}{2} \times 2\pi \Rightarrow E_b = I_n \cdot \pi \cdot dA$$

Total emissive power of area dA

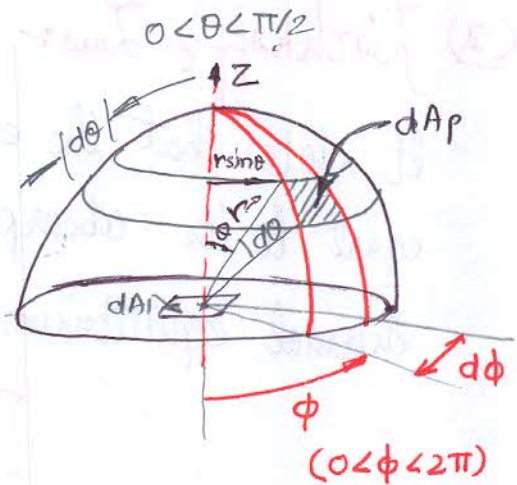
$$E_b = \sigma T^4 dA$$

$$I_n \cdot \pi \cdot dA = \sigma T^4 dA$$

$$E_b = I_n \cdot \pi$$

{ emissive power of a black body is equal to π times of intensity of radiation }

(3)



Que: 2

(a) Kirchhoff's Law:-

It states that the emissivity of the surface of a body is equal to its absorptivity when the body is in the thermal equilibrium with its surroundings.

$$\alpha = \epsilon$$

(b) Planck's Law:-

The relation for the spectral blackbody emissive power $E_{b\lambda}$ was developed by Max Planck in 1901. This relation is known as **Planck's Law** and is expressed as

$$E_{b\lambda}(\lambda, T) = \frac{C_1 \lambda^{-5}}{e^{\frac{C_2}{\lambda T}} - 1}$$

where $C_1 = 3.74177 \times 10^{-6} \text{ W-m}^2$

$C_2 = 1.438 \times 10^{-2} \text{ m-K}$.

λ = wavelength of the radiation (m)

T = Absolute temperature (K)

3 (c) :- Lambert's cosine law:-

It states that a diffuse surface radiates energy in such a manner that the rate of energy radiated in any perpendicular direction is proportional to the cosine of the angle between the direction under consideration and the normal to the surface.

$$E_{\theta} = E_{0n} \cos \theta$$

3 (d) Stefan Boltzmann law:-

It states that the energy (radiant) emitted by a body is directly proportional to the fourth power of the temperature in absolute scale. Mathematically it is written as

$$Q \propto T^4$$

$$Q = \sigma T^4$$

$$\left. \begin{array}{l} \sigma = \text{Stefan Boltzmann} \\ \text{constant} \\ = 5.67 \times 10^{-8} \text{ W/m}^2 \text{K}^4 \\ T = \text{Temp. of body (K)} \\ Q = \text{Heat flux (W/m}^2 \end{array} \right\}$$

Que: 4 Given that

$$\epsilon_1 = 0.5 = \epsilon_2$$

Let the surface temperature be T_1 and T_2 for plate 1 and 2 respectively.



* Case: 1 without shield.

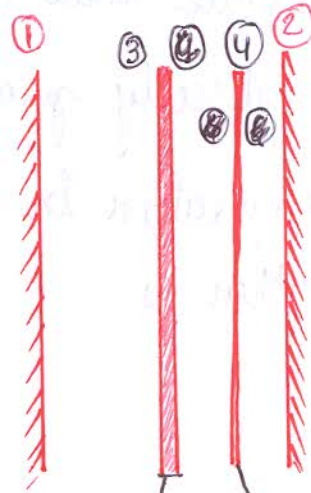
$$\left(\frac{Q_{1-2}}{A}\right)_{\text{case 1}} = \frac{\sigma(T_1^4 - T_2^4)}{\left\{\frac{1}{\epsilon_1} + \frac{1}{\epsilon_2} - 1\right\}} = \frac{\sigma(T_1^4 - T_2^4)}{3}$$

Case 1: without shield

* Case: 2

$$\frac{Q_{1-2}}{A} = \frac{\sigma(T_1^4 - T_2^4)}{\frac{1}{\epsilon_1} + \frac{1}{\epsilon_2} + 2 \sum_{i=1}^N \frac{1}{\epsilon_{3i}} - (N+1)}$$

$N = 2$ {No of shield}



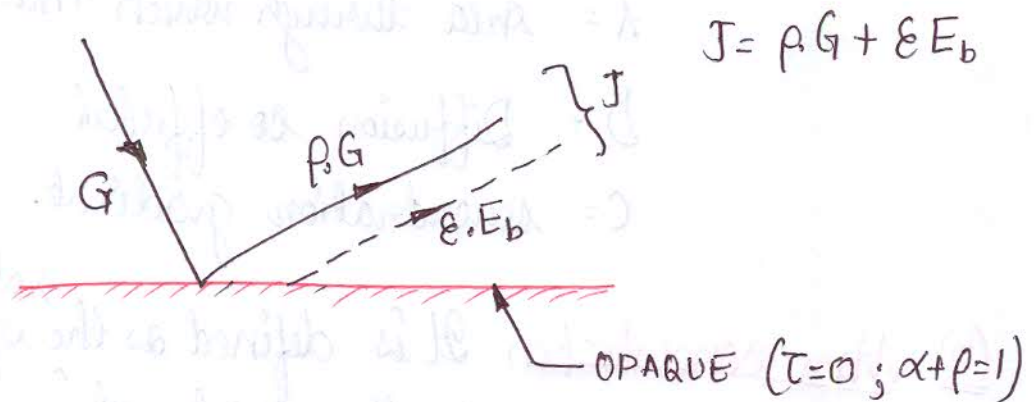
$$\left(\frac{Q_{12}}{A}\right)_{\text{case 2}} = \frac{\sigma(T_1^4 - T_2^4)}{\frac{1}{\epsilon_1} + \frac{1}{\epsilon_2} + 2 \left\{\frac{1}{\epsilon_3} + \frac{1}{\epsilon_4}\right\} - 3} = \frac{\sigma(T_1^4 - T_2^4)}{4}$$

$$\% \text{ reduction} = \frac{(Q/A)_{\text{case 1}} - (Q/A)_{\text{case 2}}}{(Q/A)_{\text{case 1}}} = \frac{92.68\%}{96.25\%}$$

Que:5

(a) Radiosity: - It is defined as the total amount of radiation leaving a surface per unit time per unit area. It is ~~given~~ written as J and expressed in W/m^2

(b) Irradiation: - It is defined as the total radiation incident upon a surface per unit time per unit area. It is expressed in W/m^2



Black Surface: - A body is said to be black, if it absorbs all incident radiation. For a black body $\alpha=1$ and $\rho=\tau=0$.

There is no such perfectly black body in nature.

Grey Surface: - A grey body has a constant monochromatic emissivity ϵ_λ with respect to wavelength. Thus grey body has a characteristic emissivity value < 1 , which does not vary with temp.

Que:6

(a) Fick's law of diffusion:-

It states that the mass flux of a constituent per unit area is proportional to the concentration gradient.

$$\frac{M_a}{A} \propto \frac{dC_A}{dx}$$

$$\frac{M_a}{A} = -D \frac{dC_A}{dx}$$

where M_A = Mass flow rate

A = Area through which mass flows

D = Diffusion coefficient

C = concentration gradient.

(b) Mass concentration It is defined as the ^{ratio} species A of a body to the total volume of mixture.

$$\text{Mass concentration} = \frac{\text{Mass concentration } A \text{ (kg/m}^3\text{)}}{\text{Volume of mixture.}}$$

Molar concentration:- It is defined as the no. of mole of species A per unit volume of mixture.

$$\text{Molar Concentration} = \frac{\text{No. of mole of species } A}{\text{Volume of mixture}} \text{ (mol/m}^3\text{)}$$

Que:7 Given that

* Temperature of surface = $540^{\circ}\text{C} + 273 = 813\text{K}$

* Black surface with 0.2m^2 area

(i) we know that the total power emitted by a body is given by Stefan Boltzmann law:

$$E_b = \sigma AT^4 = 5.67 \times 10^{-8} \times 0.2 \times (813)^4$$

$$E_b = 4954.219 \text{ W}$$

(ii) Intensity of normal radiation:-

~~The~~ $E_b = \pi \times I_n$ where $I_n =$ Normal Intensity of radiation

$$I_n = \frac{E_b}{\pi} = 15.76.97 \text{ W}$$

(iii) Maximum monochromatic wave length: The maximum monochromatic emissive power is given by the Wien's displacement law

$$\lambda_{\text{max}} \cdot T = 2.898 \times 10^{-3} \text{ m-K}$$

$$\lambda_{\text{max}} = \frac{2.898 \times 10^{-3}}{813} = 3.564 \times 10^{-6} \text{ m}$$

Que: 8 Given that

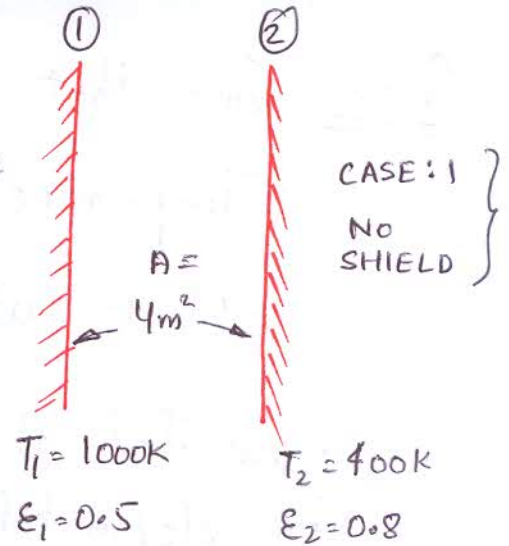
$$T_1 = 1000 \text{ K}$$

$$\epsilon_1 = 0.5$$

$$T_2 = 400 \text{ K}$$

$$\epsilon_2 = 0.8$$

$$A_1 = A_2 = 4 \text{ m}^2$$



we know that

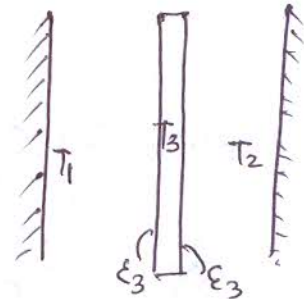
Case 1

$$(Q_{1-2})_{\text{w/o shield}} = \frac{\sigma A (T_1^4 - T_2^4)}{\frac{1}{\epsilon_1} + \frac{1}{\epsilon_2} - 1}$$

$$= \frac{5.67 \times 10^{-8} \times 4 \times (1000^4 - 400^4)}{\frac{1}{0.8} + \frac{1}{0.5} - 1} = 98219.52 \text{ W}$$

case: 2 with shield

$$N = 1$$



$$(Q_{1-2})_{\text{with shield}} = \frac{\sigma A (T_1^4 - T_2^4)}{\frac{1}{\epsilon_1} + \frac{1}{\epsilon_2} + 2 \sum_{i=1}^N \frac{1}{\epsilon_i} - (N+1)}$$

FORMULA ONLY APPLICABLE WHEN BOTH SIDE OF SHIELD HAS EQUAL ϵ

since $N = \text{no of shield} = 1$

$$Q_{1-2} = \frac{\sigma A (T_1^4 - T_2^4)}{\frac{1}{\epsilon_1} + \frac{1}{\epsilon_2} + 2 \frac{1}{\epsilon_3} - 2} \Rightarrow \frac{5.67 \times 10^{-8} \times 4 (1000^4 - 400^4)}{\frac{1}{0.8} + \frac{1}{0.5} + \frac{2}{0.4} - 2}$$

$$(Q_{1-2})_{\text{with shield}} = 35359.0272 \text{ W}$$

$$\% \text{ reduction} = \frac{98219.52 - 35359.0272}{98219.52} \times 100 = 64\%$$

(ii) Steady temperature
from energy balance

$$Q_{1-3} = Q_{3-2}$$

$$\frac{\cancel{\sigma A} (T_1^4 - T_3^4)}{\frac{1}{\epsilon_1} + \frac{1}{\epsilon_3} - 1} = \frac{\cancel{\sigma A} (T_3^4 - T_2^4)}{\frac{1}{\epsilon_3} + \frac{1}{\epsilon_2} - 1}$$

$$\frac{1000^4 - T_3^4}{\frac{1}{0.5} + \frac{1}{0.4} - 1} = \frac{(T_3^4 - 400^4)}{\frac{1}{0.4} + \frac{1}{0.8} - 1}$$

upon solving for T_3 , we get $T_3 = 821.002 \text{ K}$

— x —